

MAX-PLANCK-INSTITUT  
FÜR ASTROPHYSIK



# Physics-based preconditioning of sound waves

## A Jacobian-free Newton-Krylov method for implicit multi-D hydrodynamics

**Maxime Viallet**

Max-Planck-Institute für Astrophysik, Garching

Workshop on Newton-Krylov methods  
Lyon, October 6th, 2013

# Outline

---

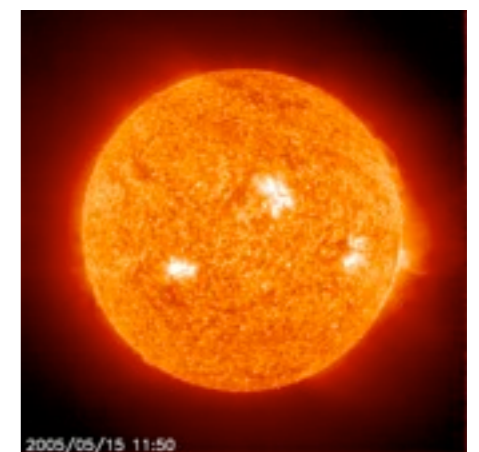
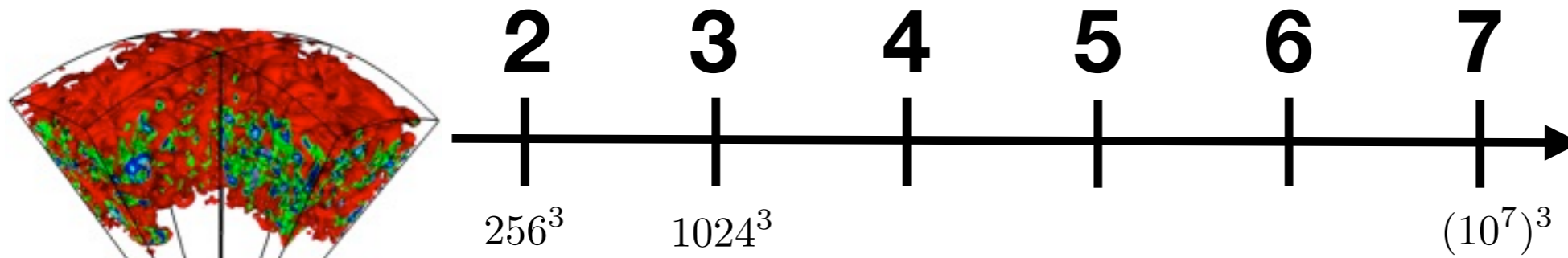
- **Introduction**
- Newton-Krylov solvers
- Conclusion

# Flows in stellar interiors

Important physical characteristics guiding modelling

---

- **Large stratification**: density decreases by orders of magnitude throughout the interior
  - **Low-Mach** flow in the interior (CZ:  $10^{-4}$ - $10^{-3}$ , RZ much smaller), but **transonic** ( $\sim 1$ ) transition toward the surface
  - Extreme parameter regime: e.g.  $Ra \sim 10^{15}$ ,  $Pr \lesssim 10^{-6}$ ,  $Re = LV/\nu > 10^{10}$
  - **Highly turbulent flow**  $\longrightarrow \frac{R}{\eta} \sim Re^{\frac{3}{4}} \sim 10^{7.5}$
- $\longrightarrow$  We will not be able to model all scales of the flow



Credit: SOHO

# Physical model

## Compressible hydrodynamics

---

- We discard low-mach approaches (e.g. anelastic approximation)
  - ➔ Modelling low-to-moderate Mach number flows in consistent background stratification
- Gas dynamics equations
  - Thermal diffusion ( $\tau \gg 1$ ) and external gravity
  - General equation-of-state (EOS)
- We do not consider (yet):
  - Rotation, multi-fluid, magnetic field
- In practice: no explicit viscosity !
  - **Implicit Large Eddy Simulation:** conservative, monotonicity-preserving schemes mimic a physical viscosity

Physical model

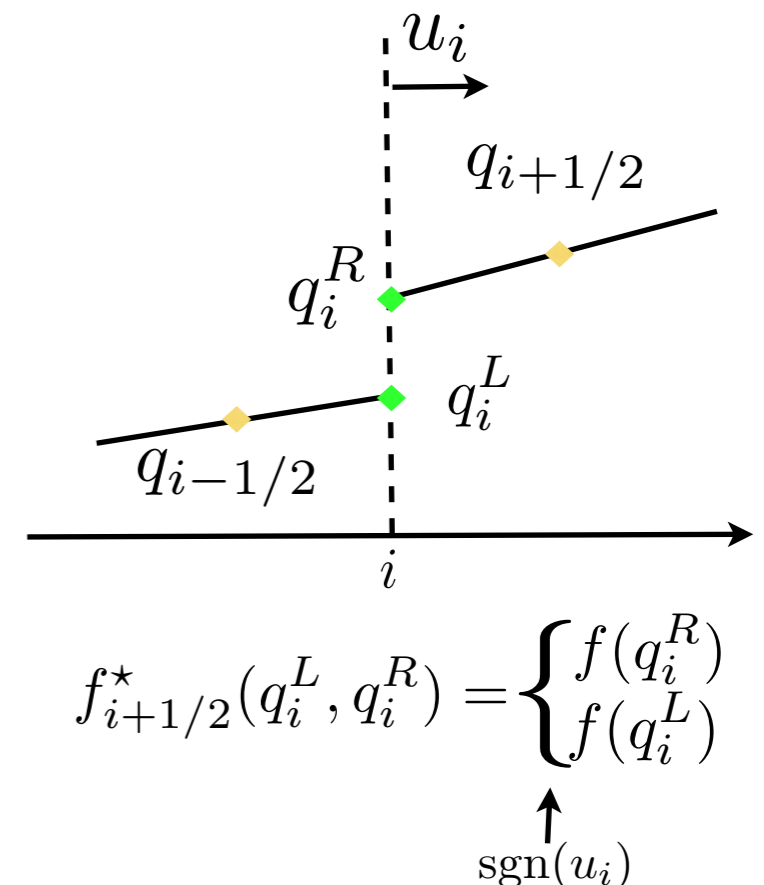
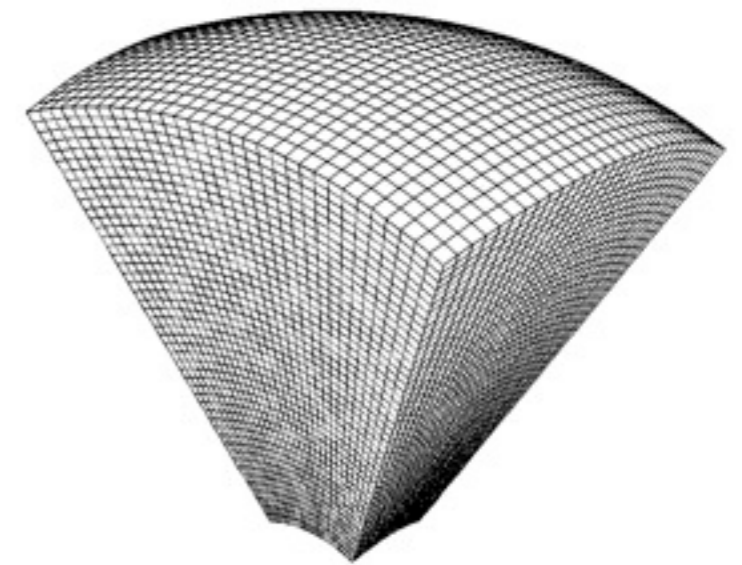
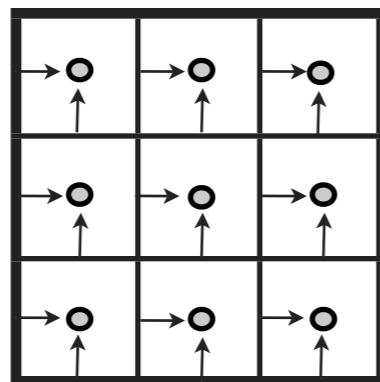
$$\begin{aligned}\partial_t \rho &= -\vec{\nabla} \cdot (\rho \vec{u}) \\ \partial_t (\rho e) &= -\vec{\nabla} \cdot (\rho e \vec{u}) - p \vec{\nabla} \cdot \vec{u} + q_{\text{visc}} + \vec{\nabla} \cdot (\chi \vec{\nabla} T) \\ \partial_t (\rho \vec{u}) &= -\vec{\nabla} \cdot (\rho \vec{u} \otimes \vec{u}) - \vec{\nabla} p + \rho \vec{g} + \vec{\nabla} \cdot \bar{\bar{\tau}}_{\text{visc}}\end{aligned}$$

$$\text{Photons conductivity: } \chi = \frac{4acT^3}{3\kappa\rho}$$

$$\text{EOS: } P, T = f(\rho, e)$$

# Spatial discretization

- Geometry:
  - Spherical coordinates (wedges) in 2D/3D
- Spatial method
  - Finite volumes on a staggered grid
  - Advection: upwind method  
“High-order donor cell” (2<sup>d</sup> order in space)
  - Diffusion: 2<sup>d</sup> order central differences



# Time integration of the equations

## Difficulties

---

- Semi-discretised system:  $\frac{dU}{dt} = R(U)$

- We have:

- **Stiff sound waves:**  $\tau_{\text{sound}} \ll \tau_{\text{phys}}$

- **Stiff diffusion:**  $\chi_r = \frac{k_r}{\rho c_p} = \frac{4acT^3}{3\kappa c_p \rho^2}$  can become very large

➔ **Severely restrict the time step when using explicit methods (CFL condition)**

$$\frac{du(t)}{dt} = f(u(t)) \longrightarrow u^{n+1} = u^n + \Delta t f(u^n)$$

➔ **Motivation for a fully implicit method**

$$\frac{du(t)}{dt} = f(u(t)) \longrightarrow u^{n+1} = u^n + \Delta t f(u^{n+1}) \quad \text{unconditionally stable (A-stability)}$$

- Time step choice is based on accuracy considerations !

➔ **MUSIC: MULTidimensional Stellar Implicit Code**

(see Viallet, Baraffe, Walder, A&A, 2011, 2013)

# Outline

---

- Introduction
- **Newton-Krylov solvers**
- Conclusion

# MUSIC

## Implicit strategy: Newton-Krylov method

---

- Fully-implicit scheme: 2<sup>d</sup> order Crank-Nicholson

- **Difficulty:** need to solve a large system of nonlinear equations  $F(U^{n+1})=0$
- 2D:  $512^2 \rightarrow N \sim 10^6$ , 3D:  $256^3 \rightarrow N \sim 84 \times 10^6$
- **Challenge:** achieve *efficiency* and *scalability*

- Strategy: Newton-Raphson method

- Initial guess:  $U^{(0)} = U^n$
- Linearization:  $J^{(k)} \delta U^{(k)} = -F(U^{(k)})$  with  $J^{(k)} \approx \frac{\partial F}{\partial U}(U^{(k)})$
- Update:  $U^{(k+1)} = U^{(k)} + \delta U^{(k)}$
- Convergence:  $\max \frac{\delta U^{(k)}}{U^{(k)}} < \epsilon \Rightarrow U^{n+1} = U^{(k+1)}$

Implicit integration using Crank-Nicholson

$$\frac{dU}{dt} = R(U) \longrightarrow$$
$$F(U^{n+1}) = U^{n+1} - U^n - \frac{\Delta t}{2} (R(U^{n+1}) + R(U^n)) = 0$$

**Most expensive step !**

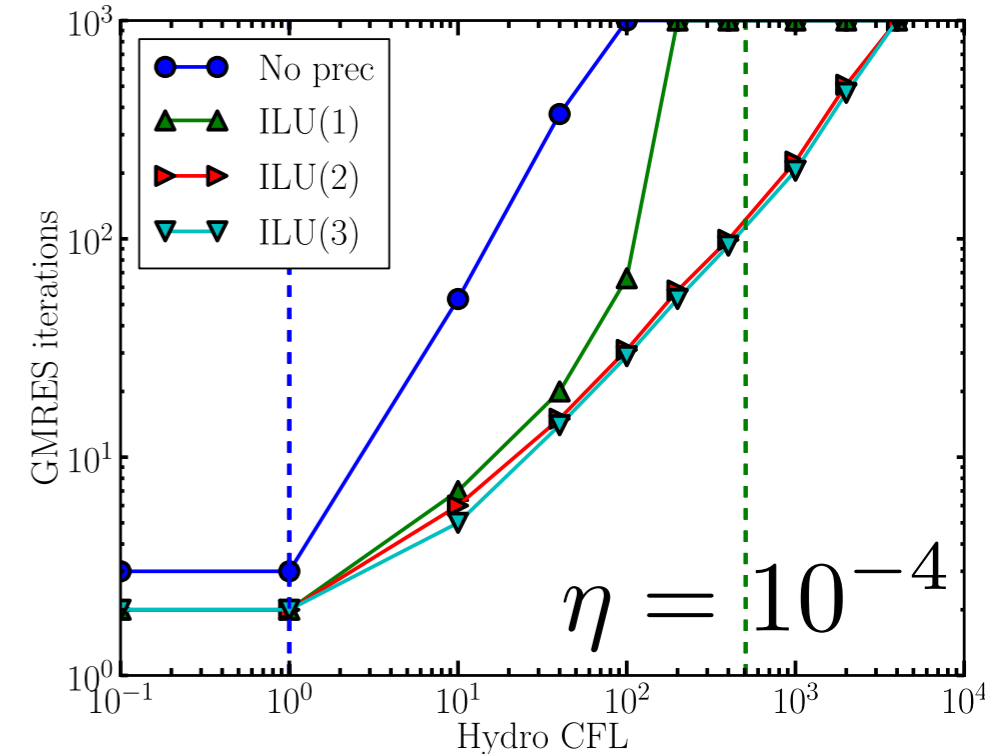
1. Build matrix J (optional !)
2. Solve sparse linear system  $Ax=b$



# MUSIC

## Newton-Krylov method

- “Inexact strategy” for  $J^{(k)}\delta U^{(k)} = -F(U^{(k)})$  which is solved with GMRES iterations (Generalized Minimum Residual Method)
  - At iteration  $p$ , construct an approximation of the solution in the Krylov space  $\mathcal{K}_p = \text{span}(F, JF, J^2F, \dots, J^{p-1}F)$
  - Accuracy can be tuned: reduce  $\|J\delta U + F\|_2$  by a factor  $\eta$  (typically loose tolerance  $\eta = 10^{-1} - 10^{-2}$ )
- Hope: a solution can be found for  $p \ll N$  (e.g.  $p \approx 30-40$ )
  - Stiffness spoils the convergence of the Krylov solver
  - **Preconditioning is necessary:** e.g. incomplete LU factorization of  $J$  (note: requires the computation of the matrix  $J$ )



Preconditioning

$$(JM^{-1})M\delta U = -F$$

Finite differences for  $J$

$$\frac{\partial F_i}{\partial U_j} \approx \frac{F(U_j + \Delta U_j) - F(U_j)}{\Delta U_j}$$

# MUSIC

## Fully implicit 3D computations

- Bottlenecks for fully-implicit 3D computations:

1. Computation/storage of the Jacobian matrix  
2D:  $\sim 50$  evaluations of  $F$ , 3D:  $\sim 100$  evaluations of  $F$
2. Incomplete LU factorization: inefficiency + cost

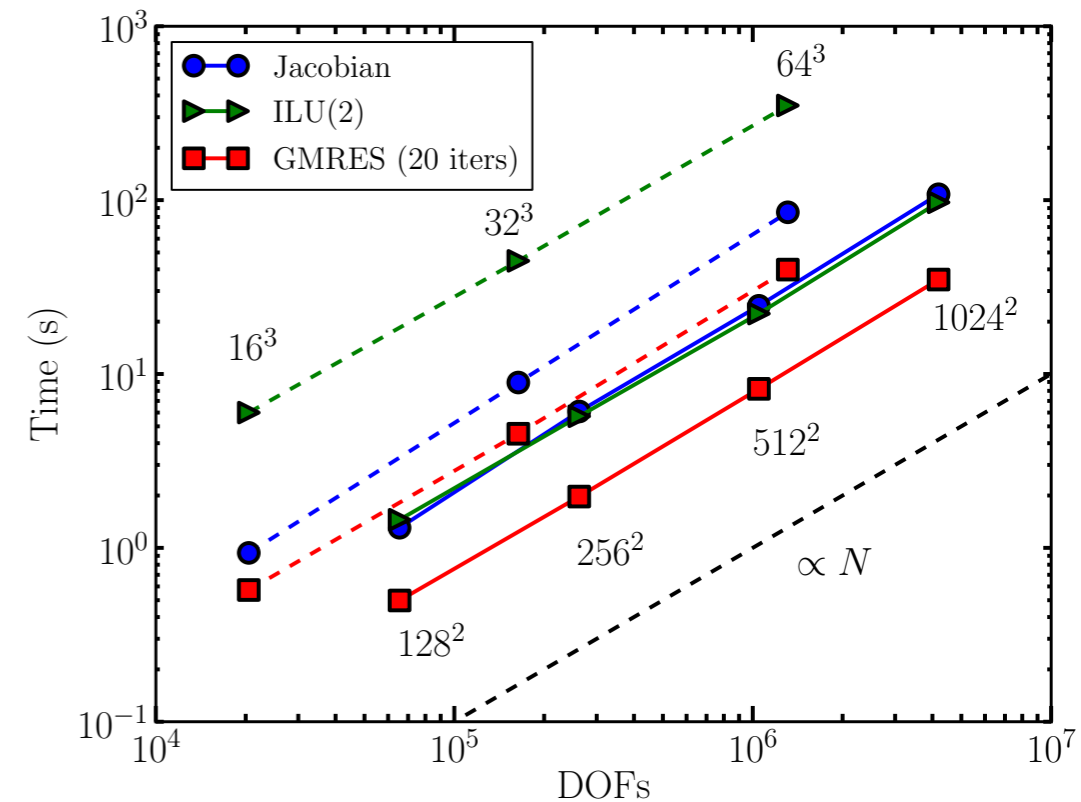
1. Use *Jacobian-free Newton-Krylov* methods  
(see review by Knoll & Keyes, JCP, 2004)

- Krylov methods do not need the Jacobian matrix, only its action on a vector

2. Physics-based preconditioning

(see e.g. Knoll et al, JSC, 25, 112, 2005, Park et al, JCP, 228, 2009)

- Origin of the problem: change of the mathematical character of acoustic fluctuations at large wave-CFL numbers



Jacobian-free approach

$$J_{\mathbf{v}} = \frac{F(\mathbf{u} + \delta \mathbf{v}) - F(\mathbf{u})}{\delta}$$

# MUSIC - 3D strategy

Jacobian-Free Newton-Krylov with physics-based preconditioning

- Semi-implicit scheme for hydrodynamics
  - Target the terms inducing sound-waves
- Solve equations for  $V = (p, e, u)$ 
  - Use implicit Euler for specific terms and explicit Euler for advection
  - Use Picard linearization for implicit terms
  - Approximate  $u^{n+1} = u^n - \Delta t \frac{1}{\rho^n} \partial_x p^{n+1}$
- Semi-implicit scheme is linear
  - Strategy: solve parabolic equation for  $\delta P$ , then get  $\delta e$  and  $\delta u$
  - Associated matrix with the “ $\delta$ -form” of the scheme

$$\begin{aligned} \partial_t p + u \partial_x p &= -\Gamma_1 p \partial_x u \\ \partial_t e + u \partial_x e &= -\frac{p}{\rho} \partial_x u \\ \partial_t u + u \partial_x u &= -\frac{1}{\rho} \partial_x p \end{aligned}$$

$$\begin{aligned} \frac{\delta p}{\Delta t} + \Gamma_1 p^n \partial_x u^{n+1} &= -u \partial_x p|^n \\ \frac{\delta e}{\Delta t} + \frac{p^n}{\rho^n} \partial_x u^{n+1} &= -u \partial_x e|^n \\ \frac{\delta u}{\Delta t} + \frac{1}{\rho^n} \partial_x p^{n+1} &= -u \partial_x u|^n \end{aligned}$$

$$\begin{aligned} \frac{\delta p}{\Delta t} - a^2 \Delta t \partial_x^2 p^{n+1} &= -u \partial_x p|^n - \Gamma_1 p^n \partial_x u^n \\ \frac{\delta e}{\Delta t} - \Delta t \frac{p^n}{(\rho^n)^2} \partial_x^2 p^{n+1} &= -u \partial_x e|^n - \frac{p^n}{\rho^n} \partial_x u^n \\ \frac{\delta u}{\Delta t} + \frac{1}{\rho^n} \partial_x p^{n+1} &= -u \partial_x u|^n \end{aligned}$$

“ $\delta$ -form”

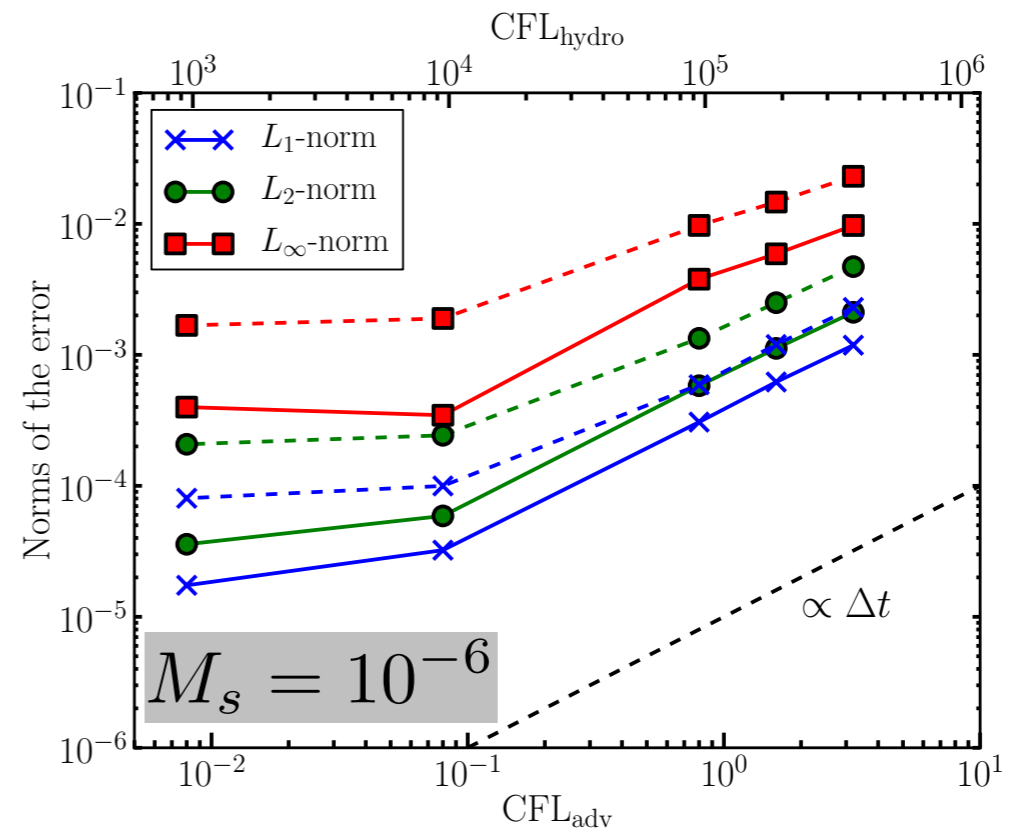
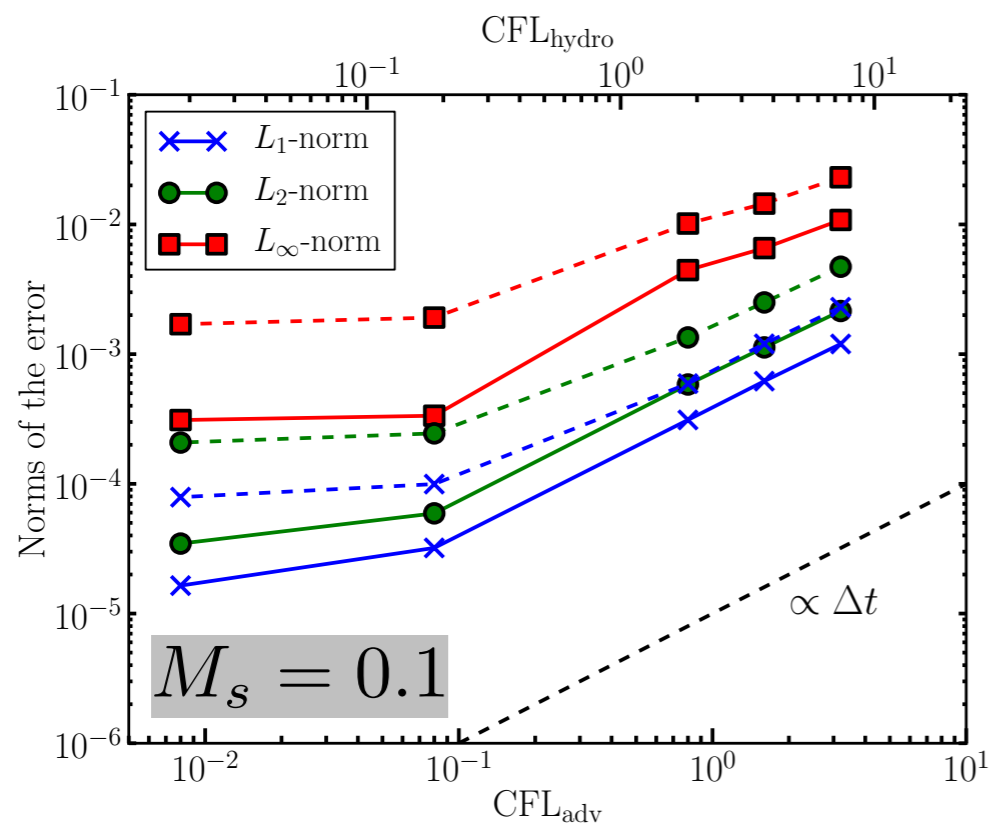
$$\tilde{J}_V \delta V = -\tilde{F}_V$$

# MUSIC - 3D strategy

Jacobian-Free Newton-Krylov with physics-based preconditioning

- Test of the semi-implicit scheme
  - Advection of an isotropic vortex in 2D
- Scheme is not prone to a wave-CFL condition
  - But advection limits the time step ( $\text{CFL}_{\text{adv}} \approx 0.2$ )

$$\text{CFL}_{\text{hydro}} = \max \left( \frac{|u| + c_s}{\Delta x} \right) \Delta t$$
$$\text{CFL}_{\text{adv}} = \max \left( \frac{|u|}{\Delta x} \right) \Delta t$$

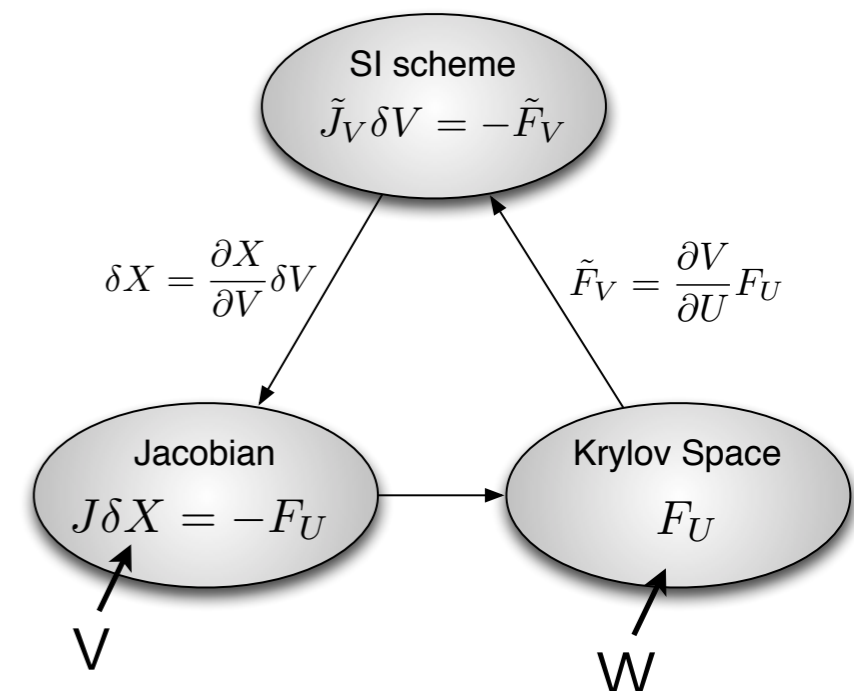


# MUSIC - 3D strategy

Jacobian-Free Newton-Krylov with physics-based preconditioning

- Three sets of variables:
  - “Conservative” variables  $U = (\rho, \rho e, \rho u)$
  - “Independent” variables  $X = (\rho, e, u)$
  - “Primitive” variables  $V = (\rho, e, u)$
- Right-Preconditioning of  $J\delta X = -F_U(X)$  with matrix  $M$ 
  - Search-space:  $\mathcal{K}_p = \text{span}(F, (JM^{-1})F, \dots, (JM^{-1})^{p-1}F)$
- GMRES algorithm: for a given Krylov vector  $w$ , compute  $(JM^{-1})w$ 
  1. Solve  $Mv = w$  for  $v$  (**Preconditioning step**)
  2. Compute  $Jv$  using a Jacobian-free approach
- PBP: interpret  $Mv=w$  as a SI step in  $\delta$ -form

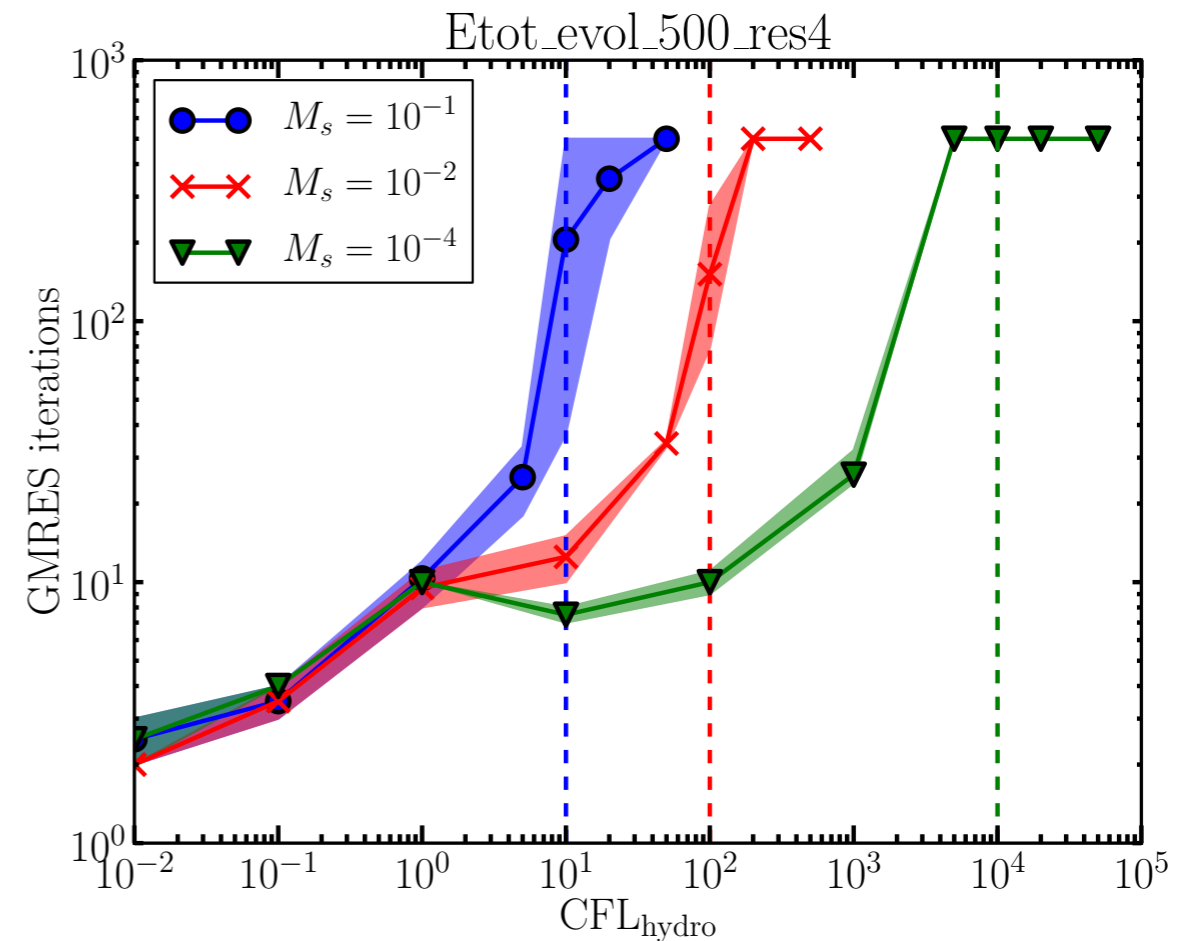
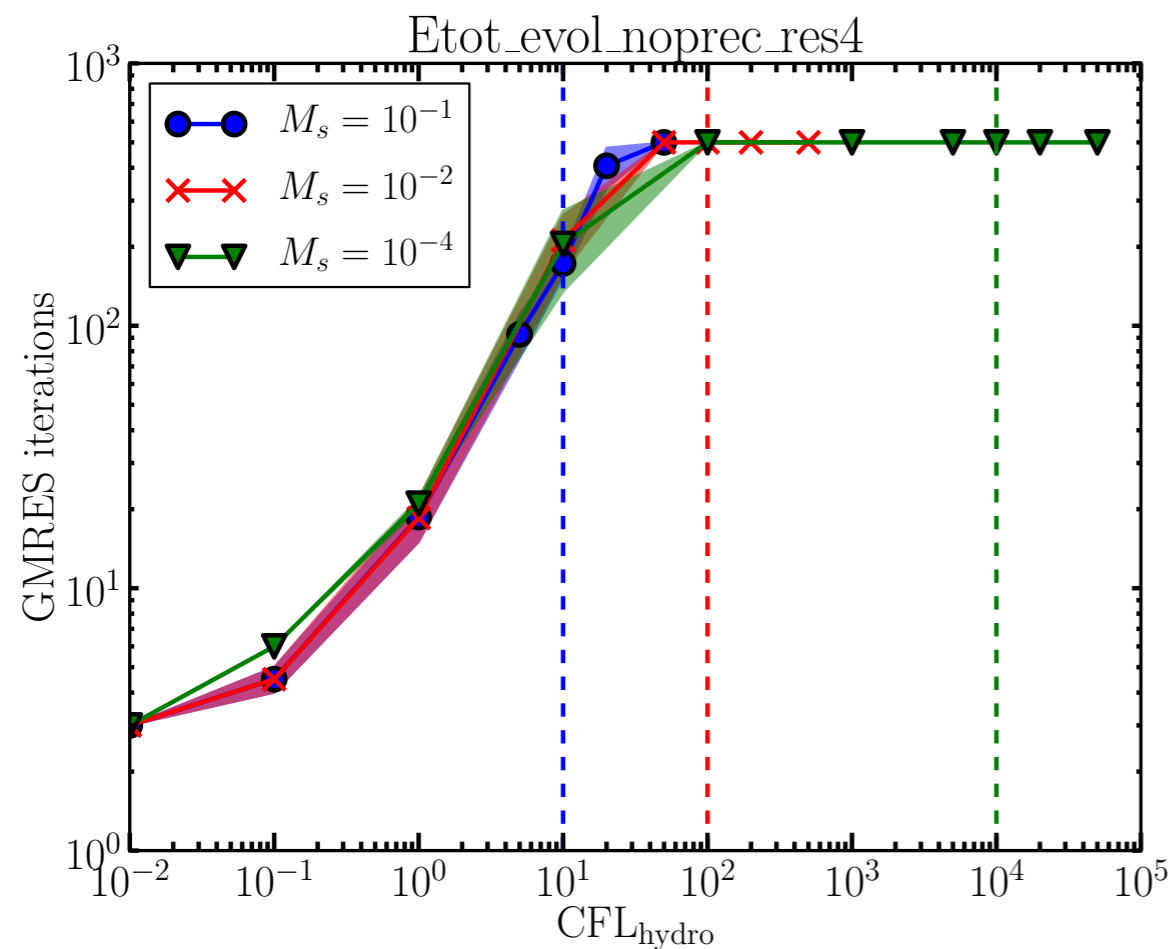
$$\begin{aligned} (JM^{-1})\delta X' &= -F_U(X) \\ M\delta X &= \delta X' \end{aligned}$$



# MUSIC - 3D strategy

Jacobian-Free Newton-Krylov with physics-based preconditioning

- Preliminary results on a 3D test: Taylor-Green vortex
  - Number iterations required to decrease the linear residual by a factor of  $\eta = 10^{-4}$



# Outline

---

- Introduction
- Newton-Krylov solvers
- **Conclusion**

# Conclusion

---

- Jacobian-Free Newton-Krylov method for hydrodynamics
  - Stiffness spoils the convergence of the Krylov solver
- The preconditioner is the most important ingredient !
  - If you know any linear method that yields an approximate solution of your (nonlinear) problem, embedding it within a JFNK method will give you the accuracy
- Unlike algebraic preconditioning, physics-based preconditioning cures the stiffness at the level of the physical model
  - Pro: obtain maximum efficiency
  - Contra: problem dependent
- Free lunch: semi-implicit schemes also provide a better initial guess for NR
- Additional physics can be included in the preconditioner
  - Thermal diffusion: results in two coupled parabolic equations for  $p$  and  $e$



END

---